Data analysis with OMC: Magnitude light curve







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Talk outline

- Standard analysis.
 - 11x11 pixel boxes.
- Extended sources.
 - Sources with low precision coordinates.
 - They are mosaics of 11x11 pixel boxes.
- Triggers.
 - To monitor GRBs almost in real time.
 - 81x81 pixel boxes (size will change to 91x91).

Required information

OMC identification number.
Example: Cyg X-1 = IOMC 2678000054
OMC Input Catalogue (fits, ascii formats)
Science Windows list.
You need a list of SCWs where your source has been observed.
ISDC Archive browser.

OMC identification I

UML - Homepage - Netscape - Retscape - Retsc	http://sdc.laeff.esa.es/omc/	
S S S S A ttp://sdc.laeff.esa.es/omc/	🖬 🔍 Search	. 🔊
2 NAC - Homepage		
ППС ППТЕБЛАL	Lavent	
Not logged in	Log in	
The OMC Archive		
This data server provides access to the INTEGRAL Optical Monitoring Camera (OM	IC) Archivo.	
The OMC Data Server includes now all Public data (with date Aug 1, 2004), PV phase and Core Pr has been re processed with the latest calibrati	ogramme OMC data up to revolution 200. The whole set of data on available.	
Resources		
Archive search and data retrieval		
System Overview		
Help Desk		
Project Documentation		
► Change your password		
The system is developed and maintained by LAEFF, based on data processed by . INTA.	ISDC. LAEFF is part of the Space Science Division of	
If you use OMC data in your research, please include the following ackn	owledgement in any resulting publications.	
Document. Done (0.372 sets)		

OMC identification II



OMC identification III



OMC identification IV

Special OMC_ID numbers:

- Observer source:
- Extension Catalogue sources:
- Dark current sub-windows:

988800000? 977700???? 99999999999

Science Windows list



Create text file with DOL list. It's useful to select SCWs by coordinates.

idx2dol index="swg_idx_omc.fits[1]" select="RA_SCX>331.5 && RA_SCX<337.7 && DEC_SCX>-7.4 && DEC_SCX<-0.2" numLog=0 outFormat=2 txtFile="OMC_5228000021.dol"

idx2dol index="swg_idx_omc.fits[1]" select="TSTART>1089 && TSTOP< 1090" numLog=0 outFormat=2 txtFile="OMC_5228000021.dol"

Terminal		
kafka:talkGeneva/beautyStars 106 >less OMC_5228000021.dol		+
/projects/omc_arc/rev_1/scw/0024/002400030010.001/swg_prp.fits[GROUPING,1,BINTAM	3LE]	
/projects/omc_arc/rev_1/scw/0024/002400030020.001/swg_prp.fits[GROUPING,1,BINTAM	3LE]	
/projects/omc_arc/rev_1/scw/0024/002400030030.001/swg_prp.fits[GROUPING,1,BINTAM	3LE _	
/projects/omc_arc/rev_1/scw/0024/002400030040.001/swg_prp.fits[GRUUPING,1,BINTAL	SLE]	1
/projects/omc_arc/rev_1/scw/0024/002400030050.001/swg_prp.fits[GRUUPING,1,BINTAL	SLE J	1
/projects/omc_arc/rev_1/scw/0024/002400030060.001/swg_prp.fits[GRUUPING,1,BINTH	SLE J	1
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/projects/omc_arc/rev_1/scw/0024/002400030000.001/swg_prp.fits[GROUPING,1,DINTHD /projects/omc_arc/rev_1/scw/0024/002400030000.001/swg_prp.fits[CROUPING,1,DINTHD		1
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/projects/omc_arc/rev_1/scw/0024/002400030170.001/swg_prp.fits[GROUPING,1,BINTAM	BLE]	11
/projects/omc_arc/rev_1/scw/0024/002400030180.001/swg_prp.fits[GROUPING,1,BINTAM	3LE]	1
/projects/omc_arc/rev_1/scw/0024/002400040010.001/swg_prp.fits[GROUPING,1,BINTAM	3LE]	
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/projects/omc_arc/rev_1/scw/0024/002400040050.001/swg_prp.fitsLGROUPING,1,BINTA	3LE J	+

Create Observation Group

Links to archive

- In -s /projects/omc_arc/rev_1/aux aux
- In -s /projects/omc_arc/rev_1/ic / ic
- In -s /projects/omc_arc/rev_1/idx idx

In -s /projects/omc_arc/rev_1/scw scw

og_create OMC_5228000021.dol instrument=OMC baseDir=. Ogid=FY-Aqr versioning=1

🔲 Terminal 💆			
kafka:obs/FY	'-Aqr.000 138 >	ll scw/	Observation group
drwxr-xr-x	3 risquez	4.0k Sep 28 19:44 002400030	directory name
drwxr-xr-x	3 risquez	4.0k Sep 28 19:44 002400030	
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drwxr-xr-x	3 risquez	4.0k Sep 28 19:44 002400030	// Jk Sep 28 19:24 IOMC 5228000021.fits
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drwxr-xr-x	3 risquez	4.0k Sep 28 19:44 002400030	/ 59 Sep 28 19:44 README
drwxr-xr-x	3 risquez	4.0K Sep 28 19:44 002400030	// Z/ Sep 28 19:43 SCM -> /projects/owc_arc/rev_1/scM/
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dowwo-wo-v	3 risquez	4 OK SEP 20 19:44 002400040	
dowwo-wo-w	3 risquez	4 OK SEP 20 19:44 002400040	
	3 risquez 3 piequez	4 OK SEP 20 19:40 002400040	
	o risquez	4,0K 3ep 20 19;40 002400040	

Execute the analysis

omc_science_analysis	No.	-
- General		<u>S</u> av
ogDOL:	/og_omc.fits[GROUPING] browse	<u>S</u> ave
startLevel:	COR -	<u>R</u> ur
endLevel:		<u>Q</u> ui
		<u>H</u> elp
-Good Time Intervals		hidde
GTI_gtiUser:	browse	
GTI_TimeFormat:	IJD 💌	
GTI_Accuracy:	any	
Course Fluxes		
- Source Fluxes -		
IMA_timestep:		
IMA_maxCentOff:	2	
IMA_numSigma:	2	
IMA_minshottime:		
IMA_maxshottime:	300 🚔	

IMA_timestep: 1 sec (don't combine shots), 630 sec (standard), 9000 sec (all data in SCW).

IMA_magboxsize: 1 pix (for sources with low flux, very high contamination, only for advanced users), 3 pix (high contamination), 5 pix (standard).

IMA_minshottime - IMA_maxshottime: Use shots between this interval. Combine shots with very different EXPOSURE times can give very noisy results for faint sources. If this variable has any value omc_science_analysis works in script mode, else it will open these selection windows.

- • ×

cd obs/FY-Aqr.000

hidden 😘

- setenv COMMONSCRIPT 1
- setenv COMMONLOGFILE omc.log
- omc_science_analysis
 ogDOL="./og_omc.fits[GROUPING]"
 startLevel=COR endLevel=IMA
 GTI_Accuracy=any IMA_timestep=1

Correction Good Time Intervals Source fluxes IC General		<u>O</u> k <u>H</u> elp
IMA_badPixels:	browse	
IMA photCal:	browse	
IMA_minSNR: 1		
IMA_skyStdDev: 10		
IMA_usePrp: ✓ checked: yes		
IMA_minTimeFrac:		
Read-out noise values specified by OMC Team		
IMA_noiseLowLeft: 45		
IMA_noiseLiowKight: 49 🚍		
IMA_noiseHighRight: 35		

Execute the analysis

Create file with final data

o_src_collect group="og_omc.fits[1]" results=omc/phot.timestep1.fits

	🔲 emacs@k	@kafka.laeff.esa.es	- • ×
	Buffers	s Files Tools Edit Search Mule Help	
Look <i>omc.log</i> file (or your <i>COMMONLOGFILE</i> variable). Scripts must finish with status=0.	Buffers Log_1	<pre>s Files Tools Edit Search Mule Help 2004-09-29716:51:36 o_src_compute_mag 4.3.2: Parameter minSNR = 1 2004-09-29716:51:36 o_src_compute_mag 4.3.2: Parameter mode = h 2004-09-29716:51:36 o_src_compute_mag 4.3.2: Parameter mode = h 2004-09-29716:51:36 o_src_compute_mag 4.3.2: Parameter mode = h 2004-09-29716:51:37 o_src_compute_mag 4.3.2: Stored MAG_V and ERRMAG_V into OMC-SRCL-RESSTAMP scw/002500010130/o es.fits[OMCSRCL-RES.1,BINTABLE] 2004-09-29716:51:38 o_src_compute_mag 4.3.2: Stored MAG_V and ERRMAG_V into OMC-SRCL-RESSTAMP scw/002500010130/o es.fits[OMCSRCL-RES.1,BINTABLE] 2004-09-29716:51:38 o_src_compute mag 4.7.2: Tack o_src_compute_mag terminating with status 0 2004-09-29716:51:38 o_src_analysis 4.8: Task o_src_analysis terminating with status 0 2004-09-29716:51:38 o_src_analysis 3.8: Pinnsmed ioou iMM oom core floxes and magnitudes 2004-09-29716:51:45 omc_science_analysis 3.8: Task omc_science_analysis terminating with status 0 2004-09-29716:51:45 omc_science_analysis 3.8: Task omc_science_analysis terminating with status 0 2004-09-29716:51:45 omc_science_analysis 3.8: task omc_science_analysis terminating with status 0 2004-09-29716:51:45 omc_science_analysis 3.8: task omc_science_analysis terminating with status 0 2004-09-30706:59:07 o_src_collect 2.2: Task o_src_collect running in SINGLE mode 2004-09-30706:59:07 o_src_collect 2.2: Parameter results = omc/phot.timestep630.fits 2004-09-30706:59:07 o_src_collect 2.2: Parameter select = 2004-09-30706:59:07 o_src_collect 2.</pre>	mc/res/omc_s\ 1,BINTABLE]
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- cd omc ۲
- fcopy "phot.timestep1.fits[OMC_ID=='5228000021']" phot.timestep1.5228000021.fits ۲

Results I

fv phot.timestep1.5228000021.fits

/•

🔲 fv: B:	nary Table of phot	t.timestep1.522800	0021.fits[1] in /pcdisk/	kafka2/risquez/talkGeneva	a/beautyStars/obs/FY-Aqr.	.000/omc/	
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1	24	002400030010	1.088355615813E+03	1.088353010988E+03	1.000000534058E+02	1.000000 <mark>5</mark> 34058E+02	2
2	24	002400030010	1.088358865236 E +03	1.088356 <mark>2</mark> 60109E+03	3.000001525879E+01	3.000001 <mark>5</mark> 25879E+01	2
3	24	002400030010	1.088361191626E+03	1.088358 <mark>5</mark> 86283 E+ 03	1.000000572204E+01	1.000000 <mark>5</mark> 72205E+01	2
4	24	002400030010	1.088361941048E+03	1.088359 <mark>335635E+03</mark>	1.000000534058E+02	1.000000 <mark>5</mark> 34058E+02	2
5	24	002400030010	1.088365193364E+03	1.088362 <mark>5</mark> 87650E+03	3.000001716614E+01	3.000001 <mark>716614E+01</mark>	² 1:Photometric
6	24	002400030010	1.088367519754E+03	1.088364913825E+03	1.000000476837E+01	1.000000 <mark>4</mark> 76837E+01	² 2:Science
7	24	002400030010	1.088368269176E+03	1.088365 <mark>663177E+</mark> 03	1.000000534058E+02	1.000000 <mark>5</mark> 34058E+02	2 2.00101100
8	24	002400030010	1.088371525833E+03	1.088368919533E+03	3.000001525879E+01	3.000001 <mark>5</mark> 25879E+01	2
9	24	002400030010	1.088373856563E+03	1.088371250048E+03	1.000000572204F+01	1 000000\$72205F+01	2
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GO TO:	E	αιτ cell:					

Results II

• fv phot.timestep1.5228000021.fits

🔲 fv: Binary 1	Table of ph	ot.timeste	p1,5228000021,fi	ts[1] in /pcdisk/	kafka2/risquez/talkGenev	a/beautyStars/obs/FY-Aqr.	.000/omc/ 3	- 0	
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5		2	5228000021	2	3.341458333300E+02	-3.813333330000E+00	1.531184806799E+02	2.889184451816E+00	1.340235276554E+02
6		3	5228000021	2	3.341458333300E 02	-3.813333330000E+00	1.499834907251E+02	6.239954096164E+00	1.290398441427E+02
		4	5228000021	2	3.341458333300E+02	-3.813333330000E+00	1.586138979049E+02	1.398527551013E+00	1.361182401126E+02
8		5	5228000021	2	3. 34145 4. Db of		1.647851507439E+02	3.072715749758E+00	1.418532097935E+02
9		6	5228000021	2	3. 34145 T.Phot		1.415041272096E+02	6.327439963802E+00	1.395649752298E+02
10		7	5228000021	2	3.34145 2:Scier		1.483848388374E+02	1.358593780495E+00	1.459994598608E+02
12		8	5228000021	2	3.34145 <mark>0333300£+02</mark>	-3.0133333300000E+00	1.487161211849E+02	2.822565734910E+00	1.397538432079E+02
13		9	5228000021	2	3.341458333300E+02	-3.813333330000 E+ 00	1.420855520369E+02	6.144897630192E+00	1.369871457435E+02
		10	5228000021	2	3.341458333300E+02	-3.813333330000 E+ 00	1.494181549529E+02	1.363334324798E+00	1.415428979066E+02
		11	5228000021	2	3.341458333300E+02	-3.813333330000E+00	1.373662306388E+02	2.725546035693E+00	1.360880547837E+02
Go to:		12	5228000021	2	3.341458333300E+02	-3.813333330000E+00	1.341937766287E+02	6.092891859819E+00	1.293308342480E+02
		13	5228000021	2	3.341458333300E+02	-3.813333330000E+00	1.450289213754E+02	1.345082301337E+00	1.407741538395E+02
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Results III

• fv phot.timestep1.5228000021.fits

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File	Edit	Tools							Help	electron/s
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2		5.514976785314E+	+00 1.2	62674464825E+02	8.113832228842E+00	1.186091325841E+01	7.733555141394E-01	5		1.418532097935E+02
3		1.429732387845E+	+01 1.1	42255959238E+02	2.159299884364E+01	1.273173767181E+01	1.5541997192892+00	5		1.395649752298E+02
4		2.078247123430E+	+00 1.3	47716548051E+02	2.893578792599E+00	1.122688859648E+01	2.344622981842E-01	5		1.459994598608E+02
5		6.026475962821E+	+00 1.8	30387094609 E +02	9.007079284350E+00	1.147730595298E+01	1.147245741956E+00	5		1.397538432079E+02
6		1.486758941604E+	+01 1.3	42688451829E+02	2.247804702446E+01	1.244724713934E+01	2.095563340367E+00	5		1.369871457435E+02
7		2.086966586883E+	+00 1.4	69806110340E+02	2.889024948073E+00			5		1.415428979066E+02
8		5.360584440979E+	+00 1.3	59704094330E+02	7.855964935075E+00	Flux used to	or calculate	5		1.360880547837E+02
9		1.421648568502E+	+01 1.5	36956723620E+02	2.151378011579E+01	MAG V It	can be 1.3 or !	5 5		1.293308342480E+02
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11		5.228202322159E+	+00 1.3	85921783579E+02	7.665806617809E+00	1.211747639085E+01	4.425107699832E-01	5		
12		1.423252461334E+	01 1.0	96029895992 E +02	2.147872279282E+01	1.229555447790E+01	1.491450435594E+00	5		
13		2.068672147011E+	+00 1.3	90060193386 E +02	2.865355240795E+00	1.112658731948E+01	2.125849493154E-01	5		
					-					
Go to		Edit c	ell:							
	/ _							/ / /		

Results IV

fv phot.timestep1.5228000021.fits



Results V

fv phot.timestep1.5228000021.fits



Results VI

- /cd omc
- /curveplot phot.timestep1.5228000021.fits "1088 1092" "13 11.5" "FY Agr with OMC-INTEGRAL data"

curveplot is a script that draw MAG_V versus TFIRST. It is not into OSA 4.X. You can use fv instead.

curveplot "phot.timestep1.5228000021.fits[EXPOSURE>50]" "1088 1092" "13 11.5" "FY Aqr with OMC-INTEGRAL data



Check results I

- Photometric shots: 10 stars, usually 1 shot per SCW, with 10 sec exposure.
- Faint photometrics: 5 stars, in science shots, but some of them are occasionally contaminated by other stars.

OK, photometric stars have constant flux.



Check results II

 Fast change in sensitivity during first revolutions. This is not corrected with flatfield.



Check results III

 Some MAG_V values are strange, around a few per 1000. We need to look at the sub-windows for cosmic rays and other problems (e.g. read-out noise).



Check results IV



Crowded field. You should use an smaller aperture (usually *FLUX_3*) Readout noise. They are horizontal groups of noisy pixels





Cosmic rays. They don't have a star profile (PSF around 1.4 pixels)

Saturated stars. Use a smaller *EXPOSURE*.



Problems column

Search updated information with: o_src_get_fluxes --help

🔲 Termin	nal The					_ 0	×
Log_1	:						•
Log_1	: PRUBLEN	1 FLAGS					
Log_1	:		_1:+				
Log_1	•	issues deemed to affect the qu	ality Vo th	ot the s	EMS" column of OMC _SDCL_DES		
		The problems are listed in the	toblo	e rrudu holow	LENS CUIVINN OF ONC, SKUL-KES,		
	•		Capie	DETOM*			
	*	They are stored in an Unsigned	Integ	er Regis	ter, any problem encountered is		
Log 1	:	logically ANDed to the existin	g regi	ster val	ue.		
Log_1	*		00-				
Log_1	*	Deconstruction of the total in	to its	only po	ossible component values reveal		
Log_1	:	the individual PROBLEMS.					
Log_1	:					1	
Log_1	*	Name	Val	ue	Meaning		
Log_1	:						
Log_1	*	UMC_PRUBLEM_NUNE	0		No problems		
Log_1	:		0	/004\			
LOg_I	÷ •	ONC PRODLEN_EATRHPOLHTED_NHG		8.3	Ne controid is available on is inaccurate.		
Log_1		AMC PROBLET_BAD_CENTROID	4 2	6-3	Rad PSF A default value was used		
$l \circ \sigma 1$	* *	OMC PROBLEM ANOMALOUS PSE	16	15-41	The PSE shane is anomalous		
	:	OMC PROBLEM LOW FLUX 1	32	6755	Flux of central pixel too low		
Log 1	*			()			
Log_1	*	OMC_PROBLEM_BADPIXEL_SKY	128	(2^7)	Bad pixel found in sky bgnd		
Log_1	*	OMC_PROBLEM_BADPIXEL_RIM_5	256	(2^8)	Bad pixel found in 5x5 rim		
Log_1	*	OMC_PROBLEM_BADPIXEL_RIM_3	512	(2^9)	Bad pixel found in 3x3 rim		
Log_1	*	OMC_PROBLEM_BADPIXEL_RIM_1	1024	(2^10)	Central pixel bad		
Log_1	*			(= 0 + =)			
Log_1	:	UMC_PROBLEM_SKY_ERROR	4096	(2^{12})	Sky error larger that accepted limit		
Log_1	*	UMC_PRUBLEM_UNKNUWN_MAG	8192	(2^{-13})	Magnitude could not be calculated		0
Log_1	:	UMC_PRUBLEM_EXIND_SRC	16384	(2-14)	Source is extended - flux not valid		
Log_1	:						+

Changing TimeStep

- og_clean ogDOL="./og_omc.fits[1]" endLevel="GTI"
- omc_science_analysis ogDOL="./og_omc.fits[GROUPING]" startLevel=COR endLevel=IMA GTI_Accuracy=any IMA_timestep=630
- o_src_collect group="og_omc.fits[1]" results=omc/phot.timestep630.fits



Triggers I

Crab Nebula, Sep/29, SCW 023900840010

Trigger images will be automatically created in OSA 4.2.

o_ima_build

inswg="/projects/omc_nrt/ops_1/scw/0239/023900840010.000/swg_prp.fits[1] outfitsname="023900840010.trigger.fits" trigger="yes" datalevel="PRP" &

fv 023900840010.trigger.fits &



ne cuit	10015							
Index	Extension	Туре	Dimension			View		
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□ 1	GROUPING	Binary	25 cols X 9 rows	Header	Hist	Plot	All	Select
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□ 3	OMCSKYIMA	Image	81 X 81	Header	lma	uge	Т	Table
□ 4	OMCSKYIMA	Image	81 X 81	Header	lma	uge	٢	Fable
□ 5	OMCSKYIMA	Image	81 X 81	Header	lma	ige	٦	Fable
□ 6	OMCSKYIMA	Image	81 X 81	Header	lma	ige	٦	Fable
□ 7	OMCSKYIMA	Image	81 X 81	Header	lma	ige	٦	Fable
8	OMCSKYIMA	Image	81 X 81	Header	lma	ige	٦	Table
9	OMCSKYIMA	Image	81 X 81	Header	lma	ae	1	Table

ds9 023900840010.trigger.2.fits 023900840010.trigger.3.fits 023900840010.trigger .4.fits 023900840010.trigger.5.fits 023900840010.trigger.6.fits 023900840010.trigger.7.fits 023900840010.trigger.8.fits 023900840010.trigger.9.fits &

Triggers II

- You can process data with IRAF. Use BIAS calculated in previous shots, and flatfield from IC.
- Problem: OMC pixels are very big.





Triggers III

 Standard analysis doesn't work with triggers, but you can obtain photometric data with *IRAF* or sextractor.

File Value WCS		02	:3900840	010.trigge	r.com	o.fits						
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kafka:triggers/sep30 7	&⇒less so	purces₊reg			
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# 2 Y_IMHGE	Ubject pos	sition alor	18 Y	[pixel]	
# 3 MHG_HPER	rixed aper	rture magnı	tude vector	[mag] [mag]	
# 4 MAGERR_APER	1 _ R TMOR	VECTOR TON	. Fixed aperiure mag.	[mag]	
# 3 CCC1F 110111 12 676 13 363	10 /87/	0.0087	0.250		
77 979 21 605	11 3835	0 0147	0.118		
24.000 5.647	14.0542	0.1281	0.505		
77.447 4.509	13.3350	0.0681	0.042		
65.000 4.000	14,9824	0.2946	0.000		
32.811 3.499	13.0448	0.0534	0.218		
40.738 2.633	12.9650	0.0500	0.102		-
26,452 3,000	13,5492	0.0823	0,498		
22,000 2,000	14.9755	0.2928	0,000		1
5,000 1,518	14.0594	0,1232	0,500		+

Mosaics I

- Same OMC_ID and different RANK number in output table of results.
- Be careful. Standard analysis gives strange results because it calculates a magnitude per sub-window.



Mosaics II

 You must build the full FoV CCD images. Be careful, because images are big and almost empty.

og_create GRO_J0852-46.partial.dol instrument=OMC baseDir=. ogid=GRO_J0852-46 versioning=1 &

omc_science_analysis ogDOL="./og_omc.fits[GROUPING]" startLevel=COR endLevel=COR chatter=1 &

o_ima_build inswg="obs/ GRO_J0852-46/scw/005800090010/swg_omc.fits[1]" outfitsname="005800090010.fits" datalevel="COR" &

🔲 fv: Summa	ry of 005800090010.fits i	n /pcdisk/kafl	ka2/risquez/talkGeneva/	/mosaic/obs/	′GRO_J0852-46.	000/ 5	
File Edit	Tools						Help
Index	Extension	Туре	Dimension		View		
0 []	Primary	Image	0	Header	Image	Table	
🗆 1	GROUPING	Binary	25 cols X 23 rows	Header	Hist Plot	All Select	ī
🗆 2	OMCSKYIMA	Image	1072 × 1028	Header	Image	Table	
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9 🗌	OMCSKYIMA	Image	1072 X 1028	Header	Image	Table	
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🗆 16	OMCSKYIMA	Image	1072 × 1028	Header	Image	Table	
🗆 17	OMCSKYIMA	Image	1072 × 1028	Header	Image	Table	
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Mosaics III

- Search mosaic coordinates with ds9 or fv, and use IRAF to cut images.
 - imcopy infile[923:977,32:86] outfile
- Use the same method as in trigger mode.
- Some interesting questions:
 - Usually the first shot is photometric.
 It has no science data.
 - Look the background. You can see the *EXPOSURE* cycle: 100, 100, 30, 100, and 10 seconds (currently 10, 50, 200 seconds).
 - Look the readout noise and its typical horizontal lines.



Mosaics IV

- Search mosaic coordinate or *fv*, and use *IRAF* to cut
 - imcopy infile[923:977,32
- Use the same method as mode.
- Some interesting question
 - Usually the first shot is p
 It has no science data.
 - Look the background. Yo the EXPOSURE cycle: 1 100, and 10 seconds (cu 50, 200 seconds).
 - Look the readout noise a typical horizontal lines.



 These are a few examples, but new OMC light curves are waiting for you...